

# A PRACTICAL STEREO SYSTEM BASED ON REGULARIZATION AND TEXTURE PROJECTION

Federico Tombari <sup>1,2</sup> and Kurt Konolige <sup>1</sup>

<sup>1</sup> Willow Garage Inc., Menlo Park, CA, U.S.A.

<sup>2</sup> DEIS - ARCES, University of Bologna, Italy

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Abstract: In this paper we investigate the suitability of stereo vision for robot manipulation tasks, which require high-fidelity real-time 3D information in the presence of motion. We compare spatial regularization methods for stereo and spacetime stereo, the latter relying on integration of information over time as well as space. In both cases we augment the scene with textured projection, to alleviate the well-known problem of noise in low-textured areas. We also propose a new spatial regularization method, *local smoothing*, that is more efficient than current methods, and produces almost equivalent results. We show that in scenes with moving objects spatial regularization methods are more accurate than spacetime stereo, while remaining computationally simpler. Finally, we propose an extension of regularization-based algorithms to the temporal domain, so to further improve the performance of regularization methods within dynamic scenes.

## 1 INTRODUCTION

As part of the Personal Robot project at Willow Garage, we are interested in building a mobile robot with manipulators for ordinary household tasks such as setting or clearing a table. An important sensing technology for object recognition and manipulation is short-range (30cm – 200cm) 3D perception. Criteria for this device include:

- Good spatial and depth resolution (1/10 degree, 1 mm).
- High speed (>10 Hz).
- Ability to deal with moving objects.
- Robust to ambient lighting conditions.
- Small size, cost, and power.

Current technologies fail on at least one of these criteria. Flash ladars (Anderson et al., 2005) lack depth and, in some cases, spatial resolution, and have non-gaussian error characteristics that are difficult to deal with. Line stripe systems (Curless and Levoy, 1995) have the requisite resolution but cannot achieve 10 Hz operation, nor deal with moving objects. Structured light systems (Salvi et al., 2004) are achieving reasonable frame rates and can sometimes incorporate motion, but still rely on expensive and high-powered projection systems, while being sensitive to ambient

illumination and object reflectance. Standard block-matching stereo, in which small areas are matched between left and right images (Konolige, 1997), fails on objects with low visual texture.

An interesting and early technology is the use of stereo with *unstructured* light (Nishihara, 1984). Unlike structured light systems with single cameras, stereo does not depend on the relative geometry of the light pattern – the pattern just lends texture to the scene. Hence the pattern and projector can be simplified, and standard stereo calibration techniques can be used to obtain accurate 3D measurements.

Even with projected texture, block-matching stereo still forces a tradeoff between the size of the match block (larger sizes have lower noise) and the precision of the stereo around depth changes (larger sizes “smear” depth boundaries). One possibility is to use smaller matching blocks, but reduce noise by using many frames with different projection patterns, thereby adding information at each pixel. This technique is known as *Spacetime Stereo* (STS) (Davis et al., 2005),(Zhang et al., 2003). It produces outstanding results on static scenes and under controlled illumination conditions, but moving objects create obvious difficulties (see Figure 1, bottom-left). While there have been a few attempts to deal with motion within a STS framework (Zhang et al., 2003), (Williams et al., 2005), the results are either compu-

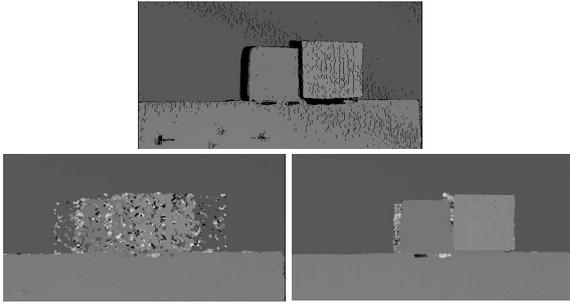


Figure 1: The top figure shows the disparity surface for a static scene; disparities were computed by integrating over 30 frames with varying projected texture using block-matching (3x3x30 block). The bottom-left figure is the same scene with motion of the center objects, integrated over 3 frames (5x5x3 block). The bottom-right figure is our local smoothing method for a single frame (5x5 block).

tationally expensive or perform poorly, especially for fast motions and depth boundaries.

In this paper, we apply regularization methods to attack the problem of motion in spacetime stereo. One contribution we propose is to enforce not only spatial, but also temporal smoothness constraints that benefit from the texture-augmented appearance of the scene. Furthermore, we propose a new regularization method, *local smoothing*, that yields an interesting efficiency-accuracy trade-off. Finally, this paper also aims at comparing STS with regularization methods, since a careful reading of the spacetime stereo literature (Davis et al., 2005; Zhang et al., 2003) shows that this has not been addressed before. Experimentally we found that, using a projected texture, regularization methods applied on single frames perform better than STS on dynamic scenes (see Figure 1) and produces interesting results also on static scenes.

In the next section we review several standard regularization methods, and introduce our novel method, *local smoothness*, which is more efficient and almost as effective. We then show how regularization can be applied across time as well as space, to help alleviate the problem of object motion in STS. In the experimental section, the considered methods are compared on static scenes and in the presence of moving objects.

## 2 SMOOTHNESS CONSTRAINTS IN STEREO MATCHING

Stereo matching is difficult in areas with low texture and at depth boundaries. Regularization methods add a smoothness constraint to model the regularity of surfaces in the real world. The general idea is to penalize those candidates lying at a differ-

ent depth from their neighbors. A standard method is to construct a disparity map giving the probability of each disparity at each pixel, and compute a global energy function for the disparity map as a multi-class Pairwise Markov Random Field. The energy is then minimized using approximate methods such as Belief Propagation (BP) (Klaus et al., 2006), (Yang, 2006) or Graph Cuts (GC) (Kolmogorov and Zabih, 2001). Even though efficient BP-based algorithms have been proposed (Yang et al., 2006), (Felzenszwalb and Huttenlocher, 2004), overall the computational load required by global approaches does not allow real-time implementation on standard PCs.

Rather than solving the full optimization problem over the disparity map, scanline methods enforce smoothness along a line of pixels. Initial approaches based on Dynamic Programming (DP) and Scanline Optimization (SO) (Scharstein and Szeliski, 2002) use only horizontal scanlines, but suffer from streaking effects. More sophisticated approaches apply SO over multiple, variably-oriented scanlines (Hirschmuller, 2005) or use multiple horizontal and vertical passes (Kim et al., 2005), (M. Bleyer, 2008), (Gong and Yang, 2005). These methods tend to be faster than global regularization, though the use of several DP or SO passes tends to increase the computational load of the algorithms.

Another limit to the applicability of these approaches within a mobile robotic platform is their fairly high memory requirements. This section we review scanline methods and proposes a new method called *local smoothness*.

### 2.1 Global Scanline Methods

Let  $I_L, I_R$  be a rectified stereo image pair sized  $M \cdot N$  and  $W(p)$  a vector of points belonging to a squared window centered on  $p$ . The *standard* block-matching stereo algorithm computes a local cost  $C(p, d)$  for each point  $p \in I_L$  and each possible correspondence at disparity  $d \in D$  on  $I_R$ :

$$C(p, d) = \sum_{q \in W(p)} e(I_L(q), I_R(\delta(q, d))). \quad (1)$$

where  $\delta(q, d)$  is the function that offsets  $q$  in  $I_R$  according to the disparity  $d$ , and  $e$  is a (dis)similarity function. A typical dissimilarity function is the  $L_1$  distance:

$$e(I_L(q), I_R(\delta(q, d))) = |I_L(q) - I_R(\delta(q, d))|. \quad (2)$$

In this case, the best disparity for point  $p$  is selected as:

$$d^* = \arg \min_d \{C(p, d)\}. \quad (3)$$

In the usual SO or DP-based framework, the global energy functional being minimized along a scanline  $S$  is:

$$E(d(\cdot)) = \sum_{p \in S} C(p, d(p)) + \sum_{p \in S} \sum_{q \in \mathcal{N}(p)} \rho(d(p), d(q)) \quad (4)$$

where  $d(\cdot)$  denotes now a function that picks out a disparity for its pixel argument, and  $q \in \mathcal{N}(p)$  are the neighbors of  $p$  according to a pre-defined criterion. Thus to minimize (4) one has to minimize two different terms, the first acting as a local evidence and the other enforcing smooth disparity variations along the scanline, resulting in a non-convex optimization problem. The smoothness term  $\rho$  is usually derived from the Potts model (Potts, 1995):

$$\rho(d(p), d(q)) = \begin{cases} 0 & d(p) = d(q) \\ \pi & d(p) \neq d(q) \end{cases} \quad (5)$$

$\pi$  being a penalty term inversely proportional to the temperature of the system. Usually for stereo a Modified Potts model is deployed, which is able to handle slanted surfaces by means of an additional penalty term  $\pi_s \ll \pi$ :

$$\rho(d(p), d(q)) = \begin{cases} 0 & d(p) = d(q) \\ \pi_s & |d(p) - d(q)| = 1 \\ \pi & \text{elsewhere} \end{cases} \quad (6)$$

Thanks to (6), smooth variations of the disparity surface are permitted at the cost of the small penalty  $\pi_s$ . Usually in SO and DP-based approaches the set of neighbours for a point  $p$  includes only the previous point along the scanline,  $p_{-1}$ . From an algorithmic perspective, an aggregated cost  $A(p, d)$  has to be computed for each  $p \in S, d \in D$ :

$$A(p, d) = C(p, d) + \min_{d'} \{A(p_{-1}, d') + \rho(d, d')\} \quad (7)$$

Because of the nature of (7) the full cost for each disparity value at the previous point  $p_{-1}$  must be stored in memory. If a single scanline is used, this typically requires  $O(M \cdot D)$  memory, while if multiple passes along non-collinear scanlines are concerned, this usually requires  $O(M \cdot N \cdot D)$  memory (Hirschmuller, 2005).

## 2.2 Local Smoothness

Keeping the full correlation surface over  $M \cdot N \cdot D$  is expensive; we seek a more local algorithm that aggregates costs incrementally. In a recent paper (Zhao and Katupitiya, 2006), a penalty term is added in a local fashion to improve post-processing of the disparity image based on left-right consistency check. Here, we apply a similar penalty during the construction of

the disparity map and generalize its use for multiple scanlines. Given a scanline  $S$ , we can modify (7) as follows:

$$A_{LS}(p, d) = C(p, d) + \rho(d, \tilde{d}) \quad (8)$$

where

$$\tilde{d} = \arg \min_d \{C(p_{-1}, d)\} \quad (9)$$

is the best disparity computed for the previous point along the scanline. Hence, each local cost is penalized if the previously computed correspondence along the scanline corresponds to a different disparity value. In this approach, there is no need to keep track of an aggregated cost array, since the aggregated cost for the current point only depends on the previously computed disparity. In practice the computation of (8) for the current disparity surface might be performed simply by subtracting  $\pi$  from  $C(p, \tilde{d})$  and  $\pi - \pi_s$  from  $C(p, \tilde{d} - 1), C(p, \tilde{d} + 1)$ .

Enforcing smoothness in just one direction helps handle low-textured surfaces, but tends to be inaccurate along depth borders, especially in the presence of negative disparity jumps. Using two scans, e.g. horizontally from left to right and from right to left, helps to reduce this effect, but suffers from the well-known streaking effect (Scharstein and Szeliski, 2002). In order to enforce inter-scanline consistency, we run local smoothness over 4 scans, 2 vertical and 2 horizontal (see Figure 8). In this case, which we will refer to as *Spatial Local Smoothness* ( $LS_s$ ), the aggregated cost (8) is modified as follows:

$$A_{LS_s}(d) = C(p, d) + \sum_{q \in \mathcal{N}(p)} \rho(d, d(q)). \quad (10)$$

Here  $\mathcal{N}$  refers to the 4 disparities previously computed on  $p$ . The computation of  $d^*$  benefits from propagated smoothness constraints from 4 different directions, which reduces noise in low-textured surfaces, and also reduces streaking and smearing effects typical of scanline-based methods.

It is worth pointing out that the  $LS_s$  approach can be implemented very efficiently by means of a two-stage algorithm. In particular, during the first stage of the algorithm, the forward-horizontal and forward-vertical passes are computed, and the result

is stored into two  $M \cdot N$  arrays. Then, during the second pass, the backward-horizontal and backward-vertical passes are processed, and within the same step the final aggregated cost (10) is also computed.

Then the best disparity is determined as in (9). Overall,

computational cost is between 3 and 4 times that of the standard local stereo algorithm. Memory requirements are also small –  $O(2 \times M \times N)$ .

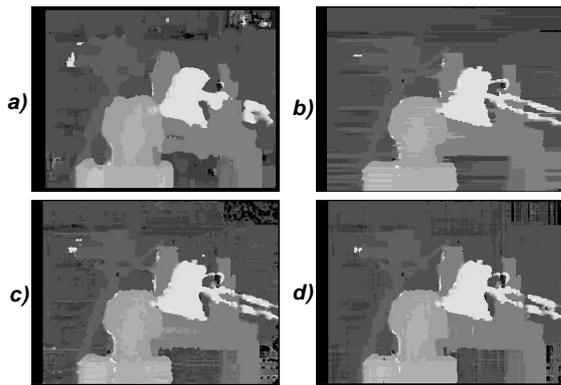


Figure 2: Qualitative comparison of different algorithms based on the smoothness constraint: a) standard b)  $SO_s$  c) local smoothness (2 horizontal scanlines) d) local smoothness (4 scanlines).

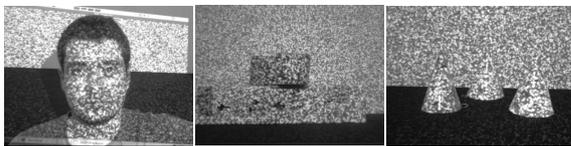


Figure 3: Dataset used for experiments: from left to right, *Face*, *Cubes*, *Cones* sequences.

### 2.3 Experimental Evaluation

In this section we briefly present some experimental results showing the capabilities of the previously introduced regularization methods on stereo data by comparing them to a standard block-correlation stereo algorithm. In particular, in addition to the  $LS_s$  algorithm, we consider a particularly efficient approach based only on one forward and one backward horizontal  $SO$  pass (M. Bleyer, 2008). This algorithm accounts for low memory requirements and fast performance, though it tends to suffer the streaking effect. We will refer to this algorithm as  $SO_s$ .

Fig. 2 shows some qualitative results on the *Tsukuba* dataset (Scharstein and Szeliski, 2002). The standard local algorithm is in (a),  $SO_s$  (b) and the  $LS_s$  algorithm in (d). Also, the figure shows the disparity map obtained by the use of the Local Smoothness criterion over only 2 horizontal scanlines in (c). It can be noticed that, compared to the standard approach, regularization methods allow for improved accuracy along depth borders. Furthermore, while methods based only on horizontal scanlines (b, c) present typical horizontal streaking effects, these are less noticeable in the  $LS_s$  algorithm (d). In our implementation, using standard incremental techniques but no SIMD or multi-thread optimization, time requirements on a standard PC for the standard,  $SO_s$  and  $LS_s$  algorithms

are 18, 62 and 65 ms, respectively.

In addition, we show some results concerning images where a pattern is projected on the scene. As for the pattern, we use a randomly-generated grayscale chessboard, which is projected using a standard video projector. Fig. 3 shows 3 frames taken from 3 stereo sequences used here and in Section 3.4 for our experiments. Sequence *Face* is a static sequence, while *Cubes* and *Cones* are dynamic scenes where the objects present in the scene rapidly shift towards one side of the table. All frames of all sequences are  $640 \times 480$  in resolution.

Figure 4 shows experimental results for the standard algorithm as well as  $SO_s$  and  $LS_s$  over different window sizes. Similarly to what done in (Davis et al., 2005), ground truth for this data is the disparity map obtained by the spacetime stereo technique (see next Section) over all frames of the sequence using a  $5 \times 5$  window patch. A point in the disparity map is considered erroneous if the absolute difference between it and the groundtruth is higher than one.

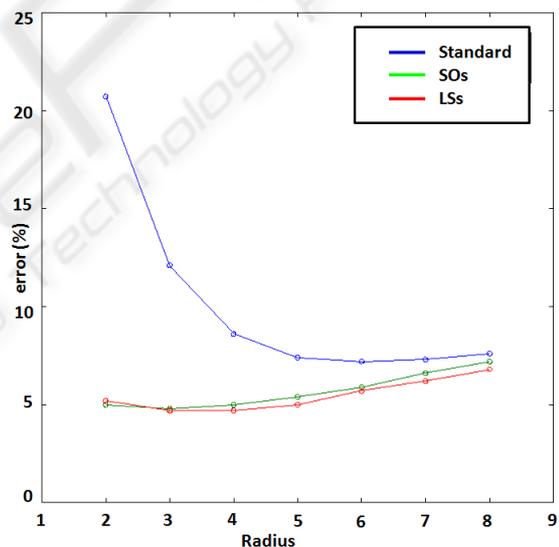


Figure 4: Quantitative comparison between different spatial approaches: standard algorithm,  $SO_s$ ,  $LS_s$ .

From the figure it is clear that, even on this real dataset, regularization methods allow for improved results compared to standard methods since the curve concerning the standard algorithm is always above the other two. It is worth pointing out that both  $SO_s$  and  $LS_s$  achieve their minimum with a smaller spatial window compared to the standard algorithm, allowing for reduced smearing effect along depth borders. Conversely, the use of regularization methods with big windows increase the error rate which tends to converge to the one yielded by the standard method. It is

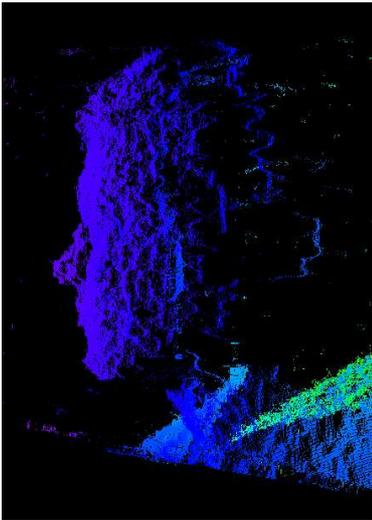


Figure 5: Point cloud showing the 3D profile of the face in Fig. 3 (left), computed using a single frame and  $LS_s$  algorithm.

also worth pointing out that overall the best result is yielded by the proposed  $LS_s$  algorithm. Finally, Figure 5 shows the 3D point cloud of the face profile obtained by using the  $LS_s$  algorithm over one frame on the *Face* dataset. From the Figure it can be noted that despite being fast and memory-efficient, this algorithm is able to obtain good accuracy in the reconstructed point cloud.

### 3 SPACETIME STEREO

Block-correlation stereo uses a spatial window to smooth out noise in stereo matching. A natural extension is to extend the window over time, that is, to use a spatio-temporal window to aggregate information at a pixel (Zhang et al., 2003), (Davis et al., 2005) (Figure 6). The intensity at position  $I(p, t)$  is now dependent on time, and the block-matching sum over a set of frames  $F$  and a spatial window  $W$  can be written as

$$C(p, d) = \sum_{t \in F} \sum_{q \in W(p)} e(I_L(q, t), I_R(\delta(q), t)). \quad (11)$$

Minimizing  $C$  over  $d$  yields an estimated disparity at the pixel  $p$ . Note that we obtain added information only if the scene illumination changes within  $F$ .

As pointed out in (Zhang et al., 2003), block matching in Equation (11) assumes that the disparity  $d$  is constant over both the local neighborhood  $W$  and the frames  $F$ . Assuming for the moment that the scene is static, by using a large temporal window  $F$

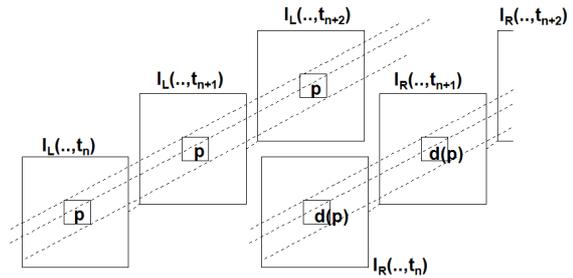


Figure 6: Spacetime window for block matching. Spatial patches centered on  $p$  are matched against corresponding patches centered on  $d(p)$ , and the results summed over all frames.

we can reduce the size of the window  $W$  while still reducing matching noise. This strategy has the further salutary effect of minimizing the smearing of object boundaries. Figure 1 (top) shows a typical result for spacetime block matching of a static scene with small spatial windows.

#### 3.1 Moving Objects

In a scene with moving objects, the assumption of constant  $d$  over  $F$  is violated. A simple scheme to deal with motion is to trade off between spatial and temporal window size (Davis et al., 2005). In this method, a temporal window of the last  $k$  frames is kept, and when a new frame is added, the oldest frame is popped off the window, and  $C(p, d)$  is calculated over the last  $k$  frames. We will refer to this approach as *sliding windows* (STS-SW). The problem is that any large image motion between frames will completely erase the effects of temporal integration, especially at object boundaries (see Figure 1, bottom-left). It is also suboptimal, since some areas of the image may be static, and would benefit from longer temporal integration.

A more complex method is to assume locally linear changes in disparity over time, that is,  $d(p, t)$  is a linear function of time (Zhang et al., 2003):

$$d(p, t) \approx d(p, t_0) + \alpha(p)(t - t_0). \quad (12)$$

For smoothly-varying temporal motion at a pixel, the linear assumption works well. Unfortunately, searching over the space of parameters  $\alpha(p)$  makes minimizing the block-match sum (11) computationally difficult. Also, the linear assumption is violated at the boundaries of moving objects, where there are abrupt changes in disparity from one frame to the next (see Figure 7). These temporal boundaries present the same kind of challenges as spatial disparity boundaries in single-frame stereo.

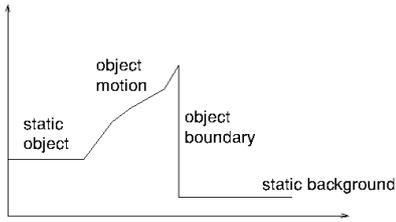


Figure 7: Disparity at a single pixel during object motion. Initially disparity is constant (no motion); then varies smoothly as the object moves past the pixel. At the object boundary there is an abrupt change of disparity.

A more sophisticated strategy would be to detect the temporal boundaries and apply temporal smoothness only up to that point. In this way, static image areas enjoy long temporal integration, while those with motion use primarily spatial information. Hence, we propose a novel method with the aim of efficiently dealing with dynamic scenes and rapidly-varying temporal boundaries. In particular, the main idea is to avoid using the spacetime stereo formulation as in (11) which blindly averages all points of the scene over time, instead enforcing a temporal smoothness constraint similarly to what is done spatially. In particular, this can be done either modelling the spatio-temporal structure with a MRF and solving using an SO or DP-based approach, or enforcing a local smoothness constraint as described in Section 2.

### 3.2 Temporal Regularization using SO

The idea of looking for temporal discontinuities was first discussed in (Williams et al., 2005), which proposed an MRF framework that extends over three frames. The problem with this approach is that the cost in storage and computation is prohibitive, even for just 3 frames. Here we propose a much more efficient method that consists in defining a scanline over time, analogous to the SO method over space. Given a cost array for each point and time instant  $C(p, d, t)$  being computed by means of any spatial method (local, global, DP-based, ...), a SO-based approach is used for propagating forward a smoothness constraint over time:

$$A_{SO}(p, d, t) = C(p, d, t) + \min_{d'} \{A_{SO}(p, d', t-1) + \rho(d, d')\} \quad (13)$$

Instead of backtracking the minimum cost path as in the typical DP algorithm, here it is more convenient to compute the best disparity over space and time as follows:

$$d^*(p, t) = \arg \min_d \{A_{SO}(p, d, t)\} \quad (14)$$

so that for each new frame its respective disparity image can be readily computed. As shown in Figure

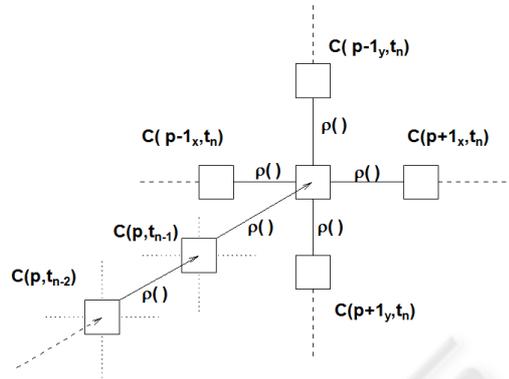


Figure 8: Local smoothing applied in the temporal domain. Disparity values influence the center pixel at time  $t_n$  from vertical and horizontal directions, and also from previous frames  $t_i, i < n$ .

8, accumulated costs from previous frames  $t_{i < n}$  are propagated forward to influence the correlation surface at time  $t_n$ . Here we propose to use as spatial algorithm the SO-based approach deploying two horizontal scanlines as discussed in Section 2. This algorithm is referred to as  $SO_{s,t}$ .

### 3.3 Temporal Regularization using Local Smoothness

In a manner similar to applying SO across frames, we can instead use local smoothness. The key idea is to modify the correlation surface at position  $p$  and time  $t$  according to the best disparity found at the same point  $p$  at the previous instant  $t-1$ . This does not require storing and propagating a cost array, only the correspondences found at the previous time instant.

The local temporal smoothness criterion is orthogonal to the strategy adopted for solving stereo over the spatial domain, hence any local or global stereo techniques can be used together with it. Here we propose to use local spatial smoothness described in Section 2. The cost function at pixel  $p$  and time  $t$  becomes:

$$A_{LS,t}(p, d, t) = C(p, d, t) + \sum_{q \in \mathcal{N}} \rho(d, d(q, t)) + \rho(d, d(p, t-1)), \quad (15)$$

That is, the penalty terms added to the local cost are those coming from the 4 independent scanline-based processes at time  $t$  plus an additional one that depends on the best disparity computed at position  $p$  at the previous time instant (see Figure 8). This algorithm will be referred to as  $LS_{s,t}$ .

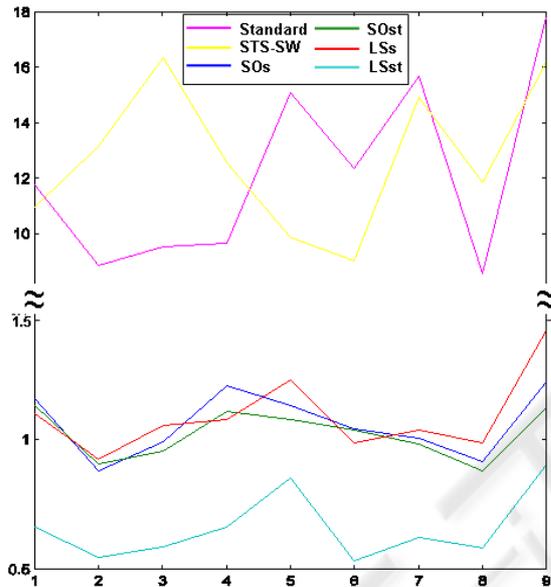
It is possible to propagate information both forwards and backwards in time, but there are several reasons for only going forwards. First, it keeps the

Table 1: Percentage of errors, *Cubes* stereo sequence.

Radius	STS-SW	Standard	$SO_s$	$SO_{s,t}$	$LS_s$	$LS_{s,t}$
2	12.8	12.1	1.1	1.0	1.1	0.7

Table 2: Percentage of errors, *Cones* stereo sequence.

Radius	STS-SW	Standard	$SO_s$	$SO_{s,t}$	$LS_s$	$LS_{s,t}$
1	46.9	49.9	5.3	5.2	14.8	12.2
3	35.4	15.9	4.2	4.1	8.2	6.9
5	31.9	9.6	4.6	4.5	7.0	6.1


Figure 9: Comparison of error percentages between different approaches for the *Cubes* sequence at each frame of the sequence. [The graph uses two different scales for better visualization].

data current – previous frames may not be useful for a realtime system. Second, the amount of computation and storage is minimal for forward propagation. Only the previous image local costs have to be maintained, which is  $O(M \cdot N)$ . In contrast, to do both forwards and backwards smoothing we would need to save local costs over  $k$  frames ( $O(k \cdot M \cdot N)$ ), and worse, recompute everything for the previous  $k$  frames, where  $k$  is the size of the temporal window for accumulation.

### 3.4 Experiments

This section presents experimental results over two stereo sequences with moving objects and a projected pattern, referred to as *Cubes* and *Cones* (see Fig. 3). To obtain ground truth for the stereo data, each differ-

ent position of the objects is captured over 30 frames with a  $3 \times 3$  spatial window, and stereo depths are averaged over time by means of spacetime stereo. Then, a sequence is built up by using only one frame for each different position of the objects.

As a comparison, we compute spacetime stereo using the sliding window approach (STS-SW). This approach is compared with regularization techniques based only on spatial smoothness (i.e.  $SO_s$ ,  $LS_s$ ) as well as with those enforcing temporal regularization (i.e.  $SO_{s,t}$ ,  $LS_{s,t}$ ).

Figure 9 shows the error rates of each algorithm for each frame of the *Cubes* dataset, with a fixed spatial window of radius 2. Table 1 reports the average error over the whole sequence. In addition, Figure 1 shows the ground truth for one frame of the sequence as well as the results obtained by STS-SW and  $LS_{s,t}$ . As can be seen, due to the rapid shift of the objects in the scene, the approach based on spacetime stereo is unable to improve the results compared to the standard algorithm. Instead, approaches based on spatial regularization yield very low error rates, close to those obtained by the use of spacetime stereo over the same scene but with no moving objects. Furthermore, Figure 9 shows that the error variance of the methods enforcing the smoothness constraint is notably lower than that reported by the standard and STS-SW algorithms. It is worth pointing out that the use of the proposed LS regularization technique both in space and time yields the best results over all the considered frames.

As in the previous experiment, Table 2 shows the mean error percentages over the *Cones* dataset with different spatial windows (i.e., radius 1, 3, and 5). Also in this case, regularization approaches achieve notably lower error rates compared to standard and spacetime approaches. From both experiments it is possible to observe that the introduction of temporal smoothness always helps improving the performance of the considered regularization methods.

## 4 CONCLUSIONS AND FUTURE WORK

In this paper we investigated the capabilities of a 3D sensor comprised of a stereo camera and a texture projector. With off-the-shelf hardware and under real illumination conditions, we have shown that in the presence of moving objects single-frame stereo with regularization produces much better results than STS. Moreover, the proposed regularization approach based on local smoothness, though not based on a global optimization, shows good performance and reduced computational requirements. Finally, we have found that the proposed introduction of temporal smoothness helps improving the performance of the considered regularization methods.

We are currently actively developing a small, low-power stereo device with texture projection. There are two tasks that need to be accomplished. First, we are trying to optimize the local smoothness constraint to be real time on standard hardware, that is, to run at about 30 Hz on 640x480 images. Second, we are designing a small, fixed pattern projector that will replace the video projector. The challenge here is to project enough light while staying eye-safe and having a compact form factor. Using the methods developed in this paper, we believe we can make a truly competent realtime 3D device for near-field applications.

The code concerning the regularization methods and the STS algorithms used in this paper is open source and available online <sup>1</sup>.

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